

Projet de Traitement du Signal : Détection de transitoires sur des signaux d'alimentation électrique d'avion

4 avril 2013

1 Introduction

Sur un avion, l'usure des gaines d'isolation des câbles d'alimentation électriques (figure 1) peut engendrer un phénomène de conduction entre phases. Cela provoque l'apparition d'arcs qui se propagent le long des câbles (phénomène d'"arc tracking"). Ce phénomène est précédé de faibles perturbations des signaux électriques (transitoires, cf. figure 2) qui doivent être détectés le plus tôt possible afin de changer les circuits avant que les équipements ne soient endommagés.

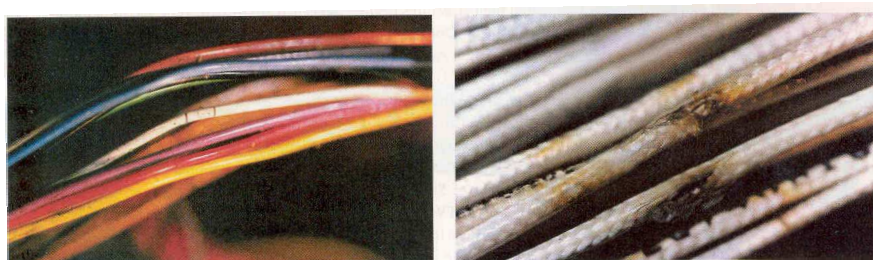


FIG. 1 – Usure des gaines d'isolation

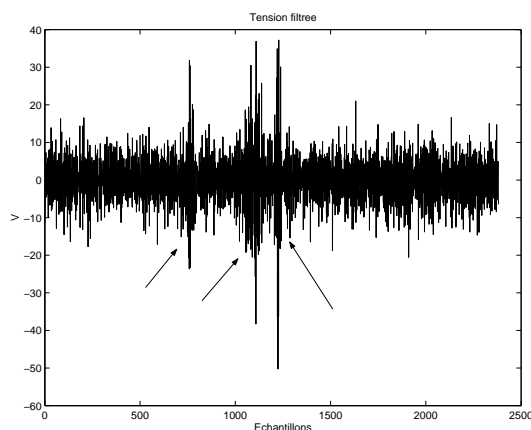


FIG. 2 – Transitoire visible sur un signal de tension filtré

Ce projet concerne la détection des transitoires sur des signaux de tension. Le réseau électrique de l'avion sera assimilé au réseau d'alimentation terrestre à 50 Hz. Le phénomène qui nous intéresse affecte les signaux à des fréquences supérieures à 500 Hz. C'est pourquoi on applique aux signaux un filtrage passe-haut afin de mettre en évidence les transitoires, qui sont très peu énergétiques par rapport au signal à 50 Hz. On procède alors à la détection de transitoires sur les signaux filtrés : pour chaque période de $\frac{1}{50} = 20$ ms, le détecteur donne une décision binaire : 1 si un transitoire est détecté, 0 sinon. La chaîne de traitement est illustrée en figure 3.

On commencera par générer un signal synthétique à partir d'hypothèses sur les signaux réels. On travaillera par la suite sur ce signal, ainsi que sur les signaux de tension réels fournis dans les fichiers *transitoire.mat* (avec transitoires) et *sig_ref.mat* (sans transitoire). L'analyse spectrale et temporelle des signaux nous permettra de nous orienter vers des méthodes de détection appropriées. On étudiera ici deux détecteurs, l'un temporel, et l'autre fréquentiel.

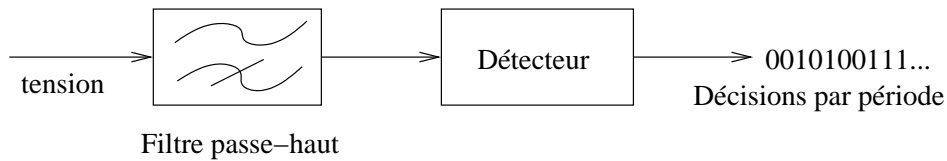


FIG. 3 – Chaîne de traitement

2 Modélisation d'un signal synthétique avec transitoire

Le réseau électrique de l'avion est alternatif, à la fréquence $f_0 = 50$ Hz. La tension varie entre -220V et +220V. La fréquence d'échantillonnage de la sonde est de $f_e = 2500$ éch/s.

1. Générer un signal sinusoïdal d'amplitude 220 (V) à la fréquence f_0 échantillonné à la fréquence f_e sur 50 périodes.
2. Ajouter à cette sinusoïde pure un bruit blanc gaussien de moyenne nulle et de variance $v_0 = 1$.
3. On modélise un transitoire par un saut de variance : ajouter à la 20^{ème} période du signal un bruit gaussien de variance $10v_0$.

3 Analyse spectrale

Déterminer le spectre du signal synthétique et du signal réel donné dans *transitoire.mat*, à l'aide des méthodes suivantes (voir annexe 1) :

1. Périodogramme
Calculer le périodogramme du signal en utilisant les fenêtres rectangulaire et de Hamming.
2. Corrélogramme
Déterminer les estimations biaisée et non biaisée de la fonction d'autocorrélation du signal (fonction *xcorr.m*). En déduire une estimation par corrélogramme de la densité spectrale de puissance du signal et comparer les résultats avec ceux de 1.

Remarque : On utilisera la fonction *load.m* pour récupérer les données du fichier.

4 Filtrage des signaux

La fréquence fondamentale (raie en f_0 dans le spectre) et ses éventuelles harmoniques masquent les transitoires qui sont beaucoup moins énergétiques. Afin de les détecter plus facilement, on

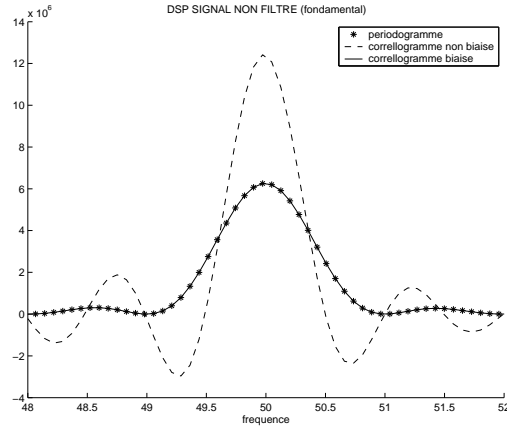


FIG. 4 – Analyse spectrale d’un signal sinusoïdal par différentes méthodes

supprime ces composantes à l’aide d’un filtre passe-haut. On trouvera en annexe 2 quelques rappels sur le filtrage analogique et numérique.

1. Calculer la réponse impulsionnelle (b_0, b_1, \dots, b_{2N}) d’un filtre non récursif passe-haut idéal, de fréquence de coupure $f_c = 500$ Hz, échantillonnée à la fréquence f_e et tronquée sur $2N + 1 = 61$ coefficients. On utilisera la méthode de la fenêtre présentée en annexe.
2. Synthétiser ce filtre, et observer sa réponse à l’aide de la fonction *freqz.m*.
3. Procéder au filtrage en utilisant la fonction *filter.m*. Que peut-on dire des N premiers points du signal filtré à l’aide de cette fonction? Afin d’accentuer l’atténuation dans la bande coupée, on filtrera le signal deux fois consécutives. On filtrera le signal synthétique, puis le signal réel (fichier *transitoire.mat*).
4. Comparer (dans les domaines temporel et fréquentiel) les signaux avant et après filtrage.

5 Un détecteur temporel : le détecteur d’énergie

On fait les hypothèses suivantes sur la nature du signal filtré en présence/absence de transitoire :

$$\begin{aligned} x_n &\sim N(0, \sigma_0^2) \text{ sous hypothèse } H_0 \text{ (pas de transitoire)} \\ x_n &\sim N(0, \sigma_1^2) \text{ sous hypothèse } H_1 \text{ (transitoire)} \end{aligned} \quad (1)$$

1. On dispose d’un signal de référence donné dans *sig_ref.mat*, dont on sait qu’il ne comporte aucun transitoire (hypothèse H_0). Observer la distribution statistique de ce signal, après filtrage passe-haut. On pourra utiliser la fonction *histfit.m* qui trace l’histogramme d’un vecteur de points ainsi que la gaussienne approchant au mieux leur distribution. Cela est-il cohérent avec (1)? Estimer la variance σ_0^2 du signal de référence filtré (fonction *var.m*).
2. Le détecteur d’énergie est le détecteur de Neyman Pearson sous les hypothèses gaussiennes exprimées ci-dessus. Calculer le rapport de vraisemblance de la v.a. (x_n) et montrer que le détecteur de Neyman Pearson peut s’exprimer ainsi :

$$H_0 \text{ rejetée si } T(x) = \sum_n x_n^2 > \gamma$$

où le seuil de détection γ dépend de la probabilité de fausse alarme P_{FA} souhaitée et de la valeur de σ_0 . On remarquera que σ_1 n’intervient pas dans le calcul du seuil. Cependant, la probabilité de détection des transitoires sera d’autant plus grande que σ_1 est élevé.

3. Quelle est la loi de la statistique de test $T(X)$ sous l'hypothèse H_0 ?
On rappelle que si (x_1, x_2, \dots, x_L) sont L variables aléatoires gaussiennes indépendantes, de moyenne nulle et de variance σ^2 , alors $\frac{1}{\sigma^2} \sum_{n=1}^L x_n^2$ est distribuée selon une loi du χ^2 à L degrés de liberté.
4. Calculer le seuil de détection γ pour une probabilité de fausse alarme fixée $P_{FA} = 10^{-15}$ en utilisant la valeur de σ_0^2 connue *a priori* (signal synthétique) ou estimée sur le signal de référence.
Remarque : Pour inverser la loi du χ^2 on pourra utiliser la fonction `chi2inv.m` (plus d'info : `help chi2inv`).
5. Appliquer l'algorithme de détection sur chaque période du signal modélisé et du signal réel.
6. Sur le signal `transitoire.mat`, procéder à l'analyse spectrale du signal filtré, sur des périodes où l'algorithme détecte. Que peut-on en déduire sur les transitoires ? Le détecteur d'énergie utilise-t-il les propriétés fréquentielles des transitoires ?
7. Sous-échantillonner d'un facteur 2 une période du signal (filtré) où un transitoire est détecté. Observer le spectre du signal sous-échantillonné. Que remarque-t-on ? Expliquer.

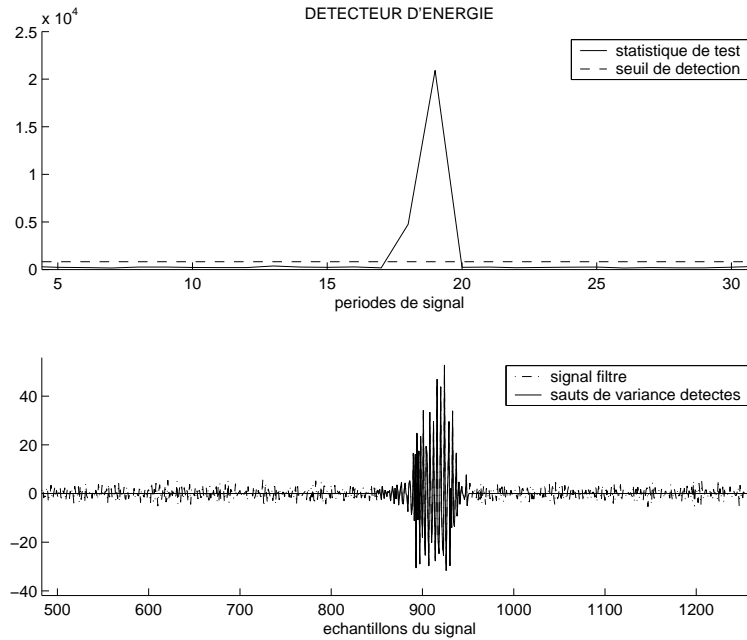


FIG. 5 – Détection d'un transitoire

6 Un détecteur fréquentiel : le détecteur de Nuttall

Le détecteur de Nuttall utilise les propriétés spectrales du transitoire à détecter. Le principe de ce détecteur est présenté dans [1], partie I. On étudiera ici le détecteur original non normalisé.

1. Montrer que lorsque $\nu = 1$, le détecteur de Nuttall est exactement équivalent au détecteur d'énergie.
2. Montrer que la loi de la statistique de test de Nuttall sous hypothèse H_0 peut être approximée par une loi normale. Les paramètres σ_0 et μ_0 de cette loi seront estimés à l'aide du signal de référence (`mean.m` pour estimer la moyenne d'un signal, `var.m` pour la variance).

3. Montrer qu'on peut exprimer la relation entre μ_0 , σ_0 , la probabilité de fausse alarme P_{FA} et le seuil de détection de Nuttall γ_{nut} par :

$$P_{FA} = Q_{N(\mu_0, \sigma_0)}(\gamma)$$

où :

$$Q_{N(\mu_0, \sigma_0)}(u) = \int_u^{+\infty} \frac{1}{\sqrt{2\pi\sigma_0^2}} e^{-\frac{(t-\mu_0)^2}{2\sigma_0^2}} dt.$$

4. Implanter le détecteur de Nuttall avec $\nu = 2$ et pour une probabilité de fausse-alarme $P_{FA} = 10^{-15}$.

Pour le calcul du seuil de détection, on pourra utiliser la fonction *norminv.m* qui inverse la fonction de répartition $f_{\mu, \sigma}$ (ou cdf : cumulative density function) de la loi normale. On rappelle que si $x \sim N(\mu, \sigma)$, alors $f_{\mu, \sigma}(a) = P(x < a) = \int_{-\infty}^a \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{(x-\mu)^2}{2\sigma^2}} dx$.

5. Comparer les résultats de détection à ceux obtenus avec le détecteur d'énergie.

Référence :

- [1] Z.Whang and P.Willett, "All-Purpose and Plug-In Power-Law Detectors for Transient Signals", *IEEE Transactions on Signal Processing*, vol. 49, n°11, nov. 2001.

ANNEXE 1 – Périodogramme / Corrélogramme

La densité spectrale de puissance (DSP) d'un signal $x(t)$ à énergie finie est définie par :

$$s(f) = TF[K_x(\tau)] = |X(f)|^2$$

où $X(f)$ est la transformée de Fourier de $x(t)$ et $K_x(\tau)$ sa fonction d'autocorrélation. Il en découle deux méthodes d'estimation de la DSP appelées **périodogramme** et **corrélogramme**.

Périodogramme : Lorsqu'on estime la transformée de Fourier avec l'algorithme de FFT rapide de Matlab, on montre qu'un estimateur satisfaisant de la DSP du signal $x(t)$ appelé périodogramme est défini par :

$$\frac{1}{N} |TFD[x(n)]|^2$$

où $x(n)$ est obtenu par échantillonnage de $x(t)$.

Corrélogramme : L'estimation de la DSP par corrélogramme comporte deux étapes :

1. Estimation de la fonction d'autocorrélation (*xcorr.m*) qui produit $\hat{K}_x(n)$.
 2. Transformée de Fourier discrète de $(\hat{K}_x(n) f(n))$, où $f(n)$ est une fenêtre de pondération et $\hat{K}_x(n)$ est l'estimation biaisée ou non biaisée de la fonction d'autocorrélation.
- *Remarque 1* : Il est important de noter que lorsque $\hat{K}_x(n)$ est l'estimateur biaisé de la fonction d'autocorrélation de $x(t)$, le corrélogramme coïncide exactement avec le périodogramme.
 - *Remarque 2* : *Estimateurs spectraux moyennés*
On montre que la variance des estimateurs de la DSP (corrélogramme et périodogramme) ne dépend pas de la durée du signal. Ces estimateurs ne sont donc pas convergents. Afin de réduire la variance des estimateurs, il est habituel de diviser le signal observé en plusieurs tranches et à moyenner les estimateurs obtenus sur chaque tranche. La variance des estimateurs moyennés est alors inversement proportionnelle au nombre de tranches, ce qui réduit la variance.
 - *Remarque 3* : *Implantation numérique*
Si le signal numérique $x(n)$ possède N_s points, la fonction *xcorr.m* calcule la fonction d'autocorrélation $\hat{K}_x(n)$ pour $n = -(N_s - 1), \dots, -1, 0, 1, \dots, (N_s - 1)$ (on a donc $2N_s - 1$ points). On peut "padding" cette autocorrélation par des zéros afin d'avoir une représentation plus précise de la DSP. L'algorithme de transformée de Fourier discrète de Matlab nécessite une symétrisation de la fonction d'autocorrélation de la façon suivante :
 - Points d'autocorrélation $\hat{K}_x(0), \hat{K}_x(1), \dots, \hat{K}_x(N_s - 1)$,
 - N_z zéros,
 - zéro central,
 - N_z zéros,
 - Points d'autocorrélation renversés $\hat{K}_x(N_s - 1), \hat{K}_x(N_s - 2), \dots, \hat{K}_x(1)$.Cette procédure de symétrisation est illustrée sur la figure 6 pour $N_s = 4$ et $N_z = 4$.

ANNEXE 2 – Du filtrage analogique au filtrage numérique

Filtrage Analogique :

Les opérations de filtrage consistent à "éliminer" ou "mettre en évidence" certaine(s) partie(s) de la bande de fréquence occupée par un signal donné. Dans le domaine continu, la transformée de Fourier d'un signal $x(t)$ est définie par $X(f) = \int_{-\infty}^{+\infty} x(t) e^{-j2\pi ft} dt$. Filtrer un signal consiste à sélectionner une bande de fréquence d'intérêt et la mettre en évidence par rapport au reste des composantes fréquentielles du signal. Dans le domaine fréquentiel (Fourier), l'opération "naturelle" de filtrage correspond donc à une multiplication du type : $X_{\text{filtre}}(f) = X(f) H(f)$, où $X(f) =$

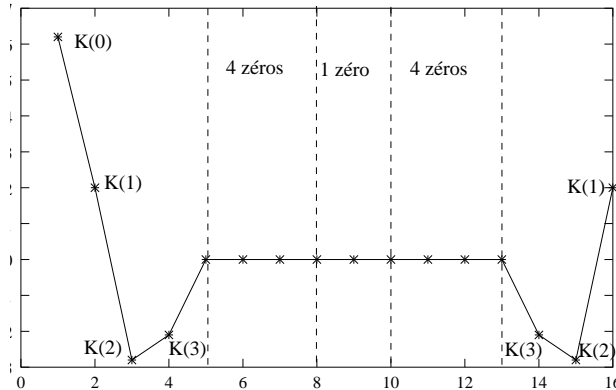


FIG. 6 – Symétrisation de la fonction d'autocorrélation

$TF \{x(t)\}$ est la transformée de Fourier du signal temporel et $H(f)$ est la **fonction de transfert** du filtre qui correspond au gabarit fréquentiel pour l'opération de filtrage désirée.

L'opération de multiplication dans le domaine fréquentiel trouve son équivalent dans la convolution dans le domaine temporel (et inversement). Ainsi une opération de filtrage temporel analogue est : $x_{\text{filtre}}(t) = x(t) * h(t)$, où $h(t)$ désigne la réponse impulsionnelle du filtre et correspond à : $h(t) = TF^{-1} \{H(f)\}$ ou encore $H(f) = TF \{h(t)\}$.

Filtrage Numérique :

Si les opérations de filtrage analogique présentées précédemment sont physiquement réalisées par des circuits électroniques (du Hardware), mettant en jeu des circuits (R, L, C), le traitement numérique mis en oeuvre par l'intermédiaire de μP et de programmation informatique connaît depuis plusieurs années un essor grandissant, offrant des possibilités beaucoup plus vastes (en termes de complexité, taille, évolutivité, reprogrammabilité).

Le signal numérique $x(n)$, $n \in \{1, \dots, N\}$ est obtenu par échantillonnage du signal continu qui représente le signal analogique associé $x(t)$. En toute rigueur, si T_e est la période d'échantillonnage, le signal numérique obtenu en sortie du convertisseur analogique-numérique est $x(n) = x(nT_e)$. Il est donc possible comme dans le cas de signaux continus de définir les opérations de filtrage dans le domaine discret. Dans cet espace de représentation, la transformée en Z est le pendant de la transformée de Fourier pour le continu :

$$X(z) = \sum_{k=-\infty}^{+\infty} x(k) z^{-k}.$$

Une des propriétés essentielle de la transformée en Z (notée TZ) est le théorème du retard :

$$X(z) z^{-m} = TZ [x(n)] z^{-m} = TZ [x(n - m)].$$

Ainsi un filtre dans le domaine discret se définit par une réponse impulsionnelle, une suite de points $h(n)$, $n \in \mathbb{N}$, dont la transformée en Z désigne la fonction de transfert du filtre :

$$H(z) = \sum_{k=-\infty}^{+\infty} h(k) z^{-k}.$$

Remarque : En pratique on n'utilise que la transformée en Z dite unilatérale, $X(z) = \sum_{k=0}^{+\infty} x(k) z^{-k}$ qui suppose le signal x causal ($x(n) = 0$ pour $n < 0$). De la même manière, le principe de causalité ("l'effet ne peut précéder la cause") fait que la réponse impulsionnelle de tout système physique est

obligatoirement nulle pour $t < 0$. Donc pour tout système causal on a : $h(n) = 0$ pour $n < 0$. C'est pourquoi nous ne noterons plus les indices de la TZ.

L'opération de filtrage dans le domaine fréquentiel, en Z, tout comme pour le cas continu est un produit de l'entrée par la fonction de transfert du filtre :

$$X_{\text{filtre}}(z) = H(z) X(z).$$

Ainsi, en utilisant le théorème du retard on montre que l'opération de filtrage peut encore s'écrire dans le domaine temporel discret :

$$X_{\text{filtre}}(z) = H(z) X(z) = \left[\sum h(k) z^{-k} \right] X(z) = \sum h(k) X(z) z^{-k}.$$

En utilisant le théorème du retard et la définition de la TZ, cette expression devient dans le domaine temporel :

$$x_{\text{filtre}}(n) = \sum_k h(k) TZ^{-1} [X(z) z^{-k}] = \sum_k h(k) x(n-k)$$

qui n'est autre que l'expression de la formule de convolution discrète. C'est l'opération de filtrage qui est concrètement implantée dans un calculateur : la sortie à un instant donné n est une combinaison linéaire des échantillons d'entrée aux instants inférieurs à n .

Cependant on vise généralement des opérations de filtrage dites temps réel (calcul de l'échantillon de sortie pendant l'intervalle de temps entre 2 échantillons d'entrée) ; or la charge de calcul de cette opération générale croît avec le temps (i.e., les indices sont tels que k varie entre 0 et l'infini dans l'équation précédente), plus on a d'échantillons passés plus le calcul d'un échantillon nécessite d'opérations élémentaires. Afin de limiter cette charge de calcul, on considère une classe de filtres linéaires ayant une charge calculatoire constante au cours du temps qui est la classe des filtres récurrents qui possèdent une expression générale de la forme :

$$x_{\text{filtre}}(n) = - \sum_{k=1}^M a_k x_{\text{filtre}}(n-k) + \sum_{k=0}^N b_k x(n-k). \quad (2)$$

Rappel : C'est cette équation temporelle qui est concrètement implantée dans le calculateur. Dans ce cas la fonction de transfert associée est une fraction rationnelle en Z :

$$H(z) = \frac{\sum_{k=0}^N b_k z^{-k}}{1 + \sum_{k=1}^M a_k z^{-k}}.$$

Pour définir un tel filtre numérique, il suffit donc de déterminer les coefficients (a_k) et (b_k) intervenant dans $H(z)$ (sous Matlab, voir la commande *filter.m*).

Filtrage non récurrent :

Lorsque les coefficients (a_k) sont nuls, le filtre est dit non récurrent. Dans ce cas, l'opération de filtrage issue de (2) est une simple convolution discrète :

$$x_{\text{filtre}}(n) = \sum_{k=0}^N b_k x(n-k).$$

Par analogie avec l'opération de filtrage analogique, on appelle alors la suite (b_k) , $k = 1, \dots, N$, la **réponse impulsionnelle du filtre**.

Synthèse d'un filtre non récursif par la méthode de la fenêtre :

Il existe une méthode simple de synthèse d'un filtre numérique non récursif : la méthode de la fenêtre consiste à calculer les coefficients (b_k) , $k = 1, \dots, N$, du filtre souhaité en échantillonnant et tronquant la réponse impulsionnelle d'un filtre analogique idéal.

Par exemple, pour synthétiser un filtre *passé-bas* sur $2N + 1$ coefficients par cette méthode, on procède de la manière suivante :

1. Expression du filtre analogique idéal :

$$\begin{aligned} H(f) &= 1 \text{ pour } f \in [-f_c, f_c], \\ &= 0 \text{ ailleurs.} \end{aligned}$$

2. Calcul de la réponse impulsionnelle du filtre analogique par transformée de Fourier inverse :

$$h(t) = TF^{-1} \{H(f)\} = \frac{\sin(2\pi f_c t)}{\pi t}.$$

3. Échantillonnage et troncature de $h(t)$: $b_0 = h\left(-\frac{N}{f_e}\right)$, ..., $b_N = h(0)$, ..., $b_{2N} = h\left(\frac{N}{f_e}\right)$.

All-Purpose and Plug-In Power-Law Detectors for Transient Signals

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Abstract—Recently, a power-law statistic operating on discrete Fourier transform (DFT) data has emerged as a basis for a remarkably robust detector of transient signals having unknown structure, location, and strength. In this paper, we offer a number of improvements to Nuttall’s original power-law detector. Specifically, the power-law detector requires that its data be prenormalized and spectrally white; a constant false-alarm rate (CFAR) and self-whitening version is developed and analyzed. Further, it is noted that transient signals tend to be contiguous both in temporal and frequency senses, and consequently, new power-law detectors in the frequency and the wavelet domains are given. The resulting detectors offer exceptional performance and are extremely easy to implement. There are no parameters to tune. They may be considered “plug-in” solutions to the transient detection problem and are “all-purpose” in that they make minimal assumptions on the structure of the transient signal, save of some degree of agglomeration of energy in time and/or frequency.

Index Terms—Crack detection, nonlinear detection, signal detection, sonar detection.

I. INTRODUCTION AND CONTEXT

A. Background

IT IS often of considerable interest to identify short-duration nonstationarities in observed signals. Applications include surveillance (e.g., [6]) in which an acoustic “transient” may indicate the presence of a threat, industrial monitoring (e.g., [16]), in which the number and severity of transients reflects machine health, and medicine (e.g., [2]). Naturally, the problem is comparatively simple if the signal to be detected is known—the only uncertainty is the time of occurrence, but knowledge of the transient is usually not available or dependable; of interest here is to detect transient signals with unknown form, location, and strength. The hypothesis test is naturally composite, with any structure open to challenge. Basically, the detector is tasked to determine whether all observations belong to a known stationary probability distribution or whether they do not.

Now, if there were nothing whatever that could be assumed about a transient signal, the detection task would be more or less hopeless. There are, fortunately, two rather qualitative properties that most transient signals possess. The first is the obvious

temporal contiguity: A transient signal is often couched as a localized burst (or bursts) in time, although the duration of such a burst is unknown in most applications. The second is a tendency for most transient signals to be bandpass, that is, it is reasonable to expect most of a transient signal’s energy to be contained in contiguous frequency observations, although again, there is usually little to be said about *which* frequencies.

To exploit only the former, and considering a transient event as a two-sided change (at some unknown time, the observations switches from having pdf f_0 to having pdf f_1 , and at a later time, there is a return to f_0), Page’s test has been explored and found to be quite useful [1], [5]. Very similar to this, Nuttall couched a transient as a contiguous burst of M bins in time, where M is known, and developed the “maximum” detector [13]. To exploit only the latter, there are detectors that begin their work on frequency domain data (usually DFT bins). Via (maximum likelihood) estimation of unknown signal parameters via the estimation–maximization (EM) algorithm, a GLRT approach is presented [21]. Of greatest interest here is Nuttall’s frequency domain “power-law” detector [12], which will be introduced shortly. It is natural to use both kinds of contiguities, and for this, we have, for example, the Gaussian-mixture time-spectrogram model in [17] and the GLRT approaches arising from linear data transformations (either time-frequency or time-scale) [3], [10], [11]. These transforms are directed toward signal representation and classification, trying to distinguish signals in the transform domain.

In [23], an attempt was made to compare the performances of a number of the above transient detection approaches on a fairly wide variety of signals. Those using time contiguity alone (Page and “maximum”) were perhaps the sturdiest performers overall but suffer from the need that certain parameters (signal strength or length) be prespecified and that data be prewhitened and prenormalized. Among the others, it was surprising that the most robust performance came from the simplest processor: Nuttall’s frequency-domain power-law statistic. It is a very good detector indeed, and in this paper, we show a number of ways to make it better still.

B. Nuttall’s Power-Law Statistic

There has been significant recent attention to Nuttall’s power-law detector [12], [14] due to its simple implementation and good performance. The test is based on the following formulation. Under the signal-absent hypothesis (H_0)—that the time-domain data is complex white Gaussian noise—preprocessing by the magnitude-square DFT yields independent and identically distributed (iid) exponential random variates. Under

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the signal-present hypothesis (H_1), the DFT observations are no longer a homogeneous population of exponentials; Nuttall's basic assumption is that there are *two* exponential populations:

$$\begin{aligned} \mathbf{H}_0: f(\mathbf{X}) &= \prod_{k=1}^N \frac{1}{\mu_0} e^{-X_k/\mu_0} u(X_k) \\ \mathbf{H}_1: f(\mathbf{X}) &= \prod_{k \notin S} \frac{1}{\mu_0} e^{-X_k/\mu_0} u(X_k) \\ &\quad \cdot \prod_{k \in S} \frac{1}{\mu_1} e^{-X_k/\mu_1} u(X_k) \end{aligned} \quad (1)$$

where

- $u(\cdot)$ unit step function (unity for positive argument and zero otherwise);
- N total number of FFT bins;
- \mathbf{X} magnitude-squared FFT bins;
- S subset with size M .

It is assumed that M signal-present bins are uniformly distributed among the N FFT bins. Clearly, the precise probability law under \mathbf{H}_1 depends on the transient signal itself, and there is no particular reason to take (1) as fact. Nevertheless, there is considerable flexibility in (1) (mostly through the unspecified S), and the detector arising from it seems to work remarkably faithfully.

At any rate, dealing with the above model, Nuttall developed power-law statistics [12] as an approximation to the optimal detector, and these have the form

$$T(\mathbf{X}) = \sum_{k=1}^N X_k^\nu \quad (2)$$

where ν is an adjustable exponent. Notice that $\nu = 1$ is the energy detector that is optimal for $M = N$, and $\nu = \infty$, which is the maximum-magnitude FFT bin, corresponds to the GLRT for $M = 1$. Through extensive computational work, it has been found that the best compromise value for ν is 2.5 when information about M is completely unavailable. Performance of this particular power-law detector is close to the best in this class of detectors. This independence from M of the power-law is fortunate.

The above power-law statistic requires prenormalized data, meaning that μ_0 in the model must be available. As an extension of power law to unknown noise level (μ_0) cases, a constant false-alarm rate (CFAR) version was introduced [15]:

$$T_{cpl}(\mathbf{X}) = \frac{\sum_{k=1}^N X_k^\nu}{\left(\sum_{k=1}^N X_k \right)^\nu}. \quad (3)$$

Clearly, T_{cpl} is not affected by a scale factor.

The statistic (2) does a yeoman's job at detecting a wide variety of block inhomogeneities, and it might be wondered why this paper, intending to improve on it, has been written. The answer is three-fold, as follows.

- 1) The statistic (2) is designed with white noise of known power in mind; the fact is that the performance of (3) is disappointing in white noise, whereas for colored noise, it has very little appeal at all. We thus extend (2) in a natural way. We estimate the noise power and normalize on a bin-by-bin basis.
- 2) The statistic (2) is essentially optimal [12] given its frequency-domain model of (1) when there is nothing whatever known about the signal-bearing set S . However, there is some tendency for real transient signals to aggregate their energy in a band, meaning that S has some structure. The challenge is to take advantage of this tendency when it exists while avoiding any degradation in performance when it does not. We believe that we have achieved this through the simple expedient of combining contiguous DFT bins.
- 3) Similar to the previous point, there is a definite tendency for real transient energy to be agglomerated in the time domain, and a magnitude-square DFT essentially destroys any such information, but there is no reason why a DFT must be the preprocessing step. We investigate the (obvious) extension that a transform other than the DFT be used.

Basically, the power-law detector is as yet neither a plug-in solution nor is it as good as it can be, and we offer some remedy here.

The organization of the paper is as follows. In Section II, we first describe the detection problem for the colored noise case and derive the associated CFAR (bin-by-bin normalized) power-law statistics. It is necessary to revisit the assumptions by which the adjustable exponent ν [see (2)] is set, and we propose a measure by which it should be chosen. In Section III, we propose extensions to exploit the contiguities of the transient and thus develop new detectors both in the frequency and the wavelet domain. It is unsatisfying to report on new detector structures without advice in threshold-setting, and in Section IV, we derive both the normal and saddlepoint approximations to the signal-absent distributions, and naturally, we compare these to simulation. Numerical comparisons between the detectors are presented in Section V, and we offer concluding remarks in Section VI.

II. CFAR POWER-LAW DETECTOR

A. Problem Description and the CFAR Power-Law Statistic

The focus of this section is to detect transients buried in colored noise with unknown but stationary spectrum. Clearly, the CFAR Power-Law in [15, eq. (3)] is to be applied to white noise and is not suitable here. As shown in Fig. 1, we write in a matrix a block of NL time domain observations as $\mathbf{x} = (\mathbf{x}_1, \mathbf{x}_2, \dots, \mathbf{x}_L)$, where \mathbf{x}_i is a column vector of dimension N whose k th element is the time sample of index $(i-1)L + k$.¹ We immediately transform each column to its magnitude-squared frequency domain equivalent \mathbf{X}_i , and record $\mathbf{X} = (\mathbf{X}_1, \mathbf{X}_2, \dots, \mathbf{X}_L)$. It is assumed that X_{j_i} s

¹To avoid possible contamination of this "reference" dataset by a transient's incipient edge, it is best to ignore a "guard" of a few blocks of data prior to that under test.

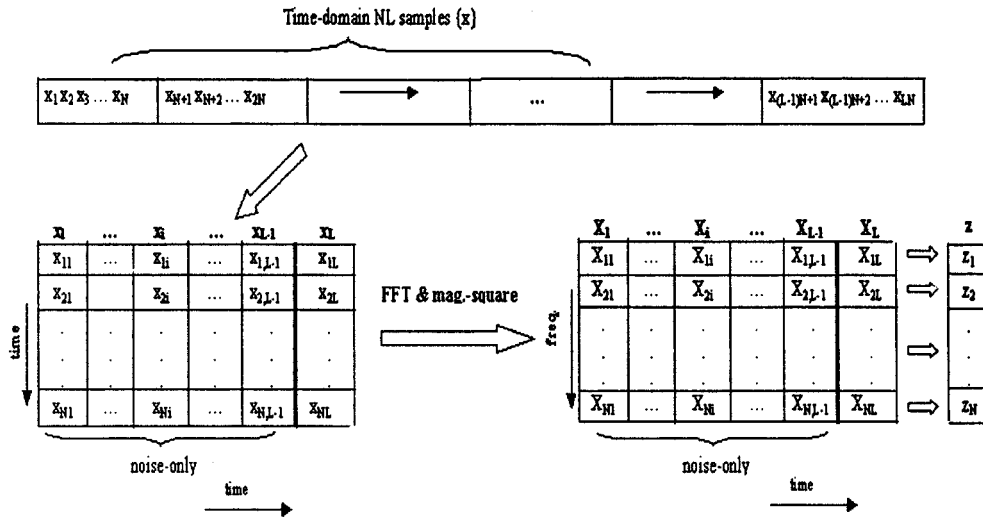


Fig. 1. Data and preliminary processing. Original time-domain sequence is reorganized into blocks, and the column-wise magnitude-square FFT is performed.

are independent and that $(\mathbf{X}_1, \dots, \mathbf{X}_{L-1})$ are known to be noise-only samples.² The probability density function (pdf) of the j th element of X_i , $i = 1, 2, \dots, L-1$ has the form

$$f(X_{ji}) = \frac{1}{\beta_j} \exp\left(\frac{-X_{ji}}{\beta_j}\right) u(X_{ji}) \quad (4)$$

where β_j are unknown but stationary.

Note that the spectral behavior of a nonwhite background is faithfully represented by the N β s. It is assumed that for the entire block of observation, the j th frequency bin maintains a mean background energy level β_j for each of the L blocks of N data. The first $L-1$ blocks provide some estimate of this level for each bin, and the goal is to test for some elevation in these levels in the *last* (L th) block. More specifically, the pdf of the X_{jL} under hypothesis H_0 follows a distribution identical to X_{ji} , $i = 1, 2, \dots, L-1$. On the other hand, when signal energy is present in the j th bin, the density of X_{jL} becomes

$$f(X_{jL}) = \frac{1}{\beta_j(1+s)} \exp\left(\frac{-X_{jL}}{\beta_j(1+s)}\right) u(X_{jL}) \quad (5)$$

where s is the relative signal power per bin, that is, the overall transient signal energy is $S_t = Ms$, in which M is the number of signal-energy-bearing DFT bins. Overall, we have the model

$$\begin{aligned} H_0: f(\mathbf{X}) &= \prod_{i=1}^L \prod_{j=1}^N \frac{1}{\beta_j} \exp\left(\frac{-X_{ji}}{\beta_j}\right) u(X_{ji}) \\ H_1: f(\mathbf{X}) &= \left[\prod_{i=1}^{L-1} \prod_{j=1}^N \frac{1}{\beta_j} \exp\left(\frac{-X_{ji}}{\beta_j}\right) u(X_{ji}) \right] \end{aligned}$$

²This assumption of independence is in practice only approximate. In what follows, for analysis, we use the assumption; our simulations are based on time domain signals, and naturally, there is a truer representation of the dependency structure.

$$\left[\prod_{j \in S} \frac{\exp\left(\frac{-X_{jL}}{\beta_j(1+s)}\right)}{\beta_j(1+s)} \prod_{j \notin S} \frac{\exp\left(\frac{-X_{jL}}{\beta_j}\right)}{\beta_j} \right] u(X_{jL}) \quad (6)$$

where S indicates the (unknown) subset with (unknown) size M out of N bins in which transient signal energy is to be found. Note that although this model may appear to have a batch flavor, Fig. 1 indicates that transient signals are to be detected on-line, although block-by-block. Each block of data that tests negative for transient signal energy joins the “window” of reference data, and hence, the least-recent block is removed to make room for it.

Following ideas similar to those frequently used in radar CFAR processing (e.g., [4]), we define the normalized magnitude-squared frequency-domain observations as

$$z_j = \frac{X_{jL}}{\frac{1}{L-1} \sum_{i=1}^{L-1} X_{ji}} \quad (7)$$

and the new power-law statistic as

$$T_{fc}(\mathbf{X}) = \sum_{j=1}^N z_j^\nu \quad (8)$$

where ν is a real exponent. Clearly, $T_{fc}(\mathbf{X})$ is non-negative and is CFAR with respect to $\{\beta_j\}$ in the model of (6). Note that normalization schemes alternative to that in (7) could be chosen. For example, one could define $z_j = X_{jL}/X_{j(R)}$ in which $X_{j(R)}$ denotes the element among $\{X_{ji}\}_{i=1}^{L-1}$ whose rank is R . It is possible that such a scheme would offer improved robustness [4], but due to its similarity to (8), we do not discuss it here.

B. SNR Analysis to Choose ν

The best value for the power ν in (8) is, in general, strongly dependent on M , which is the number of signal-present bins.

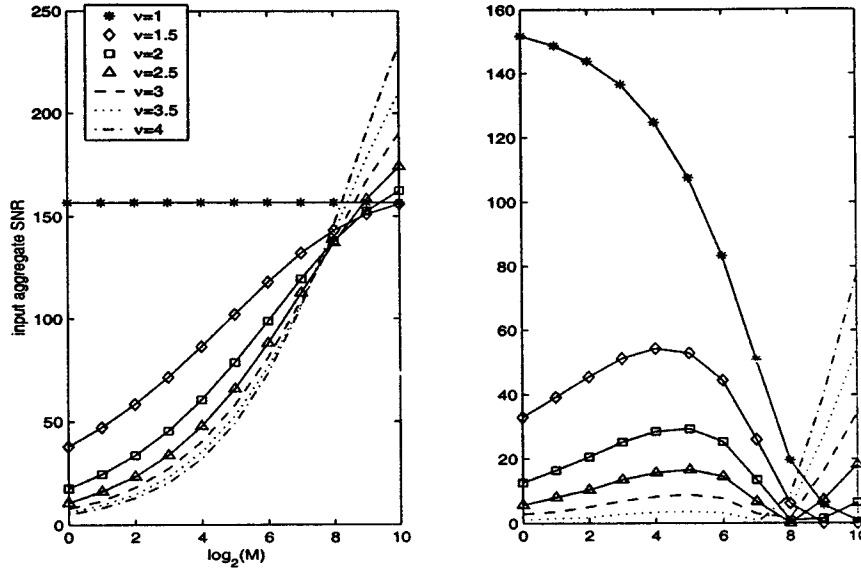


Fig. 2. SNR analysis for Nuttall's power-law statistics, with settings output $SNR = 24$, $N = 1024$. (Left) Required input aggregate SNR for power-law detectors with different ν . (Right) Input SNR-loss for different ν .

This is not at all desirable since our goal is to find a detection structure that does not depend on knowledge of such signal qualities. A clever contribution of [12] was the so-called "low-quality operating point analysis," and it was found that $2 < \nu < 3$ is a good choice over a wide range of M . This analysis depended on explicit numerical calculation of performance, and although such analysis was possible for the exponential random variables in (2), and indeed *would* be an option in (8), in some of the later statistics, it is not. Thus, here, we investigate a similar SNR analysis to suggest the best choice of ν in (8) when information about M is unavailable. Signal-to-noise ratio (SNR), which is sometimes known as deflection [19], is not a completely accurate determinant of detection performance but is a widely accepted alternative to exhaustive simulation or numerical integration. Given a statistic T , the output SNR can be expressed as

$$SNR_T = \frac{(E(T|\mathbf{H}_1) - E(T|\mathbf{H}_0))^2}{\text{Var}(T|\mathbf{H}_0)} \quad (9)$$

where $E(\cdot)$ denotes the conditional expectation, and $\text{var}(\cdot)$ is the conditional variance.

First, we exploit the SNR analysis to evaluate ν for the power-law statistics in [12]. The coincidence between our results and Nuttall's suggests that the SNR analysis is a reasonable method to choose the power ν for our new CFAR statistics. Based on statistic (2) and model (1), the associated SNR can be shown to be

$$SNR_{pl} = \frac{(M\Gamma(\nu+1)[(1+S_t/M)^\nu - 1])^2}{N[\Gamma(2\nu+1) - \Gamma(\nu+1)^2]} \quad (10)$$

where $\Gamma(\cdot)$ represents the Gamma function, and S_t denotes the total signal power in M bins, which is also referred to as the input aggregate SNR. Our purpose is to evaluate the required S_t to yield fixed output SNR_{pl} for each ν . Example results from SNR analysis for the power-law detectors with different ν are shown in Fig. 2, where $N = 1024$, and the output $SNR = 24$.

The right plot in Fig. 2 shows the input SNR-loss (ISL), which, with fixed output SNR N and M is defined as

$$ISL(\nu, M) = S_t(\nu, M) - \min_{\nu} \{S_t(\nu, M)\}. \quad (11)$$

The ISL measures the input aggregate SNR that is sacrificed through use of a fixed exponent ν , as compared with the best possible exponent ν for that M or the corresponding optimal statistic. At any rate, it is immediately seen that the best value of ν , achieving minimum average signal power per bin, changes with M and sweeps through all intermediate values. When M is completely unknown, we obtain the best compromise value for ν via $\nu = \min_{\nu} \{\max_M \{ISL\}\}$; from the figure, we find that $\nu = 2.5$ is that choice. The tendencies and results coincide well with those obtained from the low-quality operation point analysis in [12]. Based on this, we claim, as in [12], that $2 < \nu < 3$ is a good choice for a wide range of M .

Encouraged by the above, we apply the input SNR loss analysis to the detector in (8) to select ν . It is straightforward to derive the pdf of z_j under H_0 and H_1 [8], and hence, we get, after some algebra, the result

$$SNR_{fc} = \frac{M^2(L-1)(\beta(L-\nu-1, \nu+1)[(1+S_t/M)^\nu - 1])^2}{N[\beta(L-2\nu-1, 2\nu+1) - (L-1)\beta(L-\nu-1, \nu+1)^2]} \quad (12)$$

where $\beta(\cdot, \cdot)$ denotes the Beta function and S_t represents the total signal power in M bins, and $L-1$ blocks of previous DFT outputs are used for normalization. Example results for power values $\nu = 1, 1.5, 2, 2.5$, and 3 , are shown in Fig. 3, where $N = 256$, output $SNR = 6$, and $L = 10$.³ It is clear that there is no reason to explore $\nu < 1$. It is noted that the best value of ν ,

³The larger the window size L , the better the normalization, but the more susceptible the detector to a nonstationary background. The range $6 \leq L \leq 32$ is often discussed [4] and provides reasonable results; we choose $L = 10$ here as representative of that range.

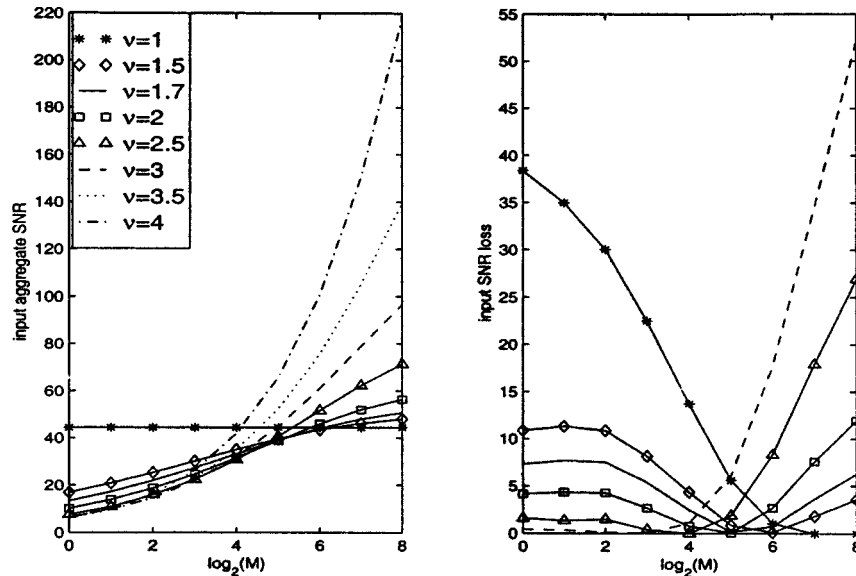


Fig. 3. SNR for CFAR power-law statistics, with settings the output $SNR = 6$, $N = 256$. Left figure: SNR for different ν ; right figure: the input SNR loss for different ν .

providing the minimum average signal power per bin with given output SNR, changes with M . Similar results will be observed by setting different N and output SNR: $1.5 < \nu < 2$ is a good choice when information of M is completely unknown, as it yields the least ISL over M . Here, $\nu = 1.7$ appears to be the best choice.

III. CONTIGUITY-BASED DETECTORS

The knowledge of signal contiguity may aid in detection. Both time and frequency contiguities are exploited in both the white- (prenormalized) and colored-noise (self-normalizing) cases to improve the detection performance. For each case, the model and the corresponding statistics are described. The detector is actually a combination of a linear and a power-law processor. Since precise contiguity information is unavailable, only the cases of two and three adjacent bins are studied in this paper. Theoretical justification for these detectors is not offered; however, they make intuitive sense, and they work well. Further, although the exponent ν could be chosen differently for difference numbers of aggregated bins, we have found that this is not a major concern, and we choose ν from the single-bin analyses of the previous section.

A. Contiguity-Based Detectors in the Frequency Domain

Just as with the power-law detectors in (2), only magnitude-squared FFT outputs are of concern here. Using the contiguity tendency in frequency, we modify Nuttall's assumption that the M signal-present bins are uniformly distributed amongst the record of N to an assumption that there is a tendency that some of the M signal-occupied bins are adjacent.

1) *Prenormalized Case*: New random variables are obtained by combining two contiguous frequency bins. We define $U_j = X_{j-1} + X_j$, $j = 1, \dots, N$. Assuming that the original $\{X_j\}$ are independent and exponential (that is, that this is Nuttall's model in which data are assumed already to have been normalized and whitened), U_j yields a $Gamma(2, \mu_0)$ random variate.

We define our new power-law detectors

$$T_{f2}(\mathbf{U}) = \sum_{j=1}^N U_j^\nu = \sum_{j=1}^N (X_{j-1} + X_j)^\nu \quad (13)$$

where $\{X_j\}$ and N have same meanings as in (1). The statistic of (13) is easily extended as

$$T_{f3}(\mathbf{U}) = \sum_{j=1}^N U_j^\nu = \sum_{j=1}^N (X_{j-2} + X_{j-1} + X_j)^\nu \quad (14)$$

to the case of three contiguous bins, and further extension is straightforward. Bins are indexed modulo $-N$.

2) *Self-Normalizing Case*: A similar combining process was adopted in the colored noise case by letting $U_{ji} = X_{j-1,i} + X_{j,i}$. This combining approach results in modified model and generates a new CFAR power-law detector in the frequency domain as

$$T_{fc2}(\mathbf{U}) = \sum_{j=1}^N \left(\frac{U_{jL}}{\frac{1}{L-1} \sum_{i=1}^{L-1} U_{ji}} \right)^\nu. \quad (15)$$

The similar detector T_{fc3} combines three contiguous bins.

B. Detectors in the Wavelet Domain

For time-domain observations, the DFT transforms a pure "time description" into a pure "frequency description" and, thus, clearly cannot take advantage of time contiguity. The discrete-time wavelet transform (DWT), which is an alternative to the DFT, is much more local and finds a good compromise—a time–frequency description. Hence, detectors in the wavelet domain will benefit from both temporal and frequency contiguity tendencies. The original work of Nuttall explored only the case the that preprocessing transformation was the

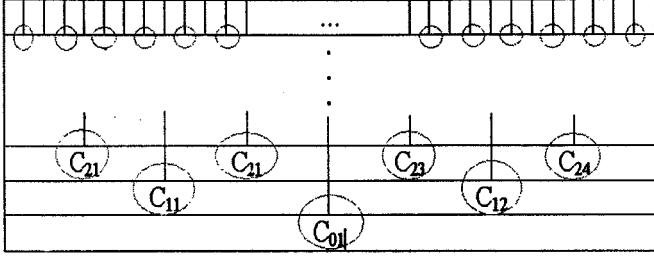


Fig. 4. Structure used in T_{w3} and T_{wc3} to combine three adjacent bins in the wavelet space. Circles illustrate the definitions $U_{kj} = C_{kj} + C_{k+1,2j-1} + C_{k+1,2j}$.

DFT. The extension to other transforms, especially the wavelet transform, is both natural and (mostly) straightforward. There are many different choices of wavelet family, and each has its proponents. However, only the simple Haar wavelet is explored due to its easy implementation, its orthogonality [20], and due to the fact that a statistic that assumes as little as possible about the transient to be detected is preferable.

1) *Prenormalized Case*: A derivation similar to that of Nuttall results in the power-law detector in the wavelet domain, considering that under a complex Gaussian noise assumption the magnitude-squared (orthonormal) DWT of the noise-only data obeys an iid exponential distribution. That is, we have

$$T_w(\mathbf{C}) = \sum_{k=0}^{K-1} \sum_{j=1}^{2^k} C_{kj}^\nu \quad (16)$$

where $K = \log_2(N)$, C_{kj} is the j th magnitude-squared DWT coefficient of k scale. The argument is that if time observations \mathbf{x} follow an iid normal distribution, the corresponding DWT vector has identical pdf [22].

In the case of preprocessing by the DFT, it was argued that there is often a tendency for transient energy to crowd into contiguous frequency bins. There is similarly a tendency for transient energy to be in nearby wavelet coefficients, as each refers to a scale that roughly matches that of the transient. As shown in Fig. 4, it is natural to adopt a tree structure in the WT case, and similar to T_{f3} of (14), we define $U_{kj} = C_{kj} + C_{k+1,2j-1} + C_{k+1,2j}$, $k = 0, 1, \dots, K-1$ and $j = 1, \dots, 2^k$ and invoke

$$T_{w3}(\mathbf{U}) = \sum_{k=0}^{K-1} \sum_{j=1}^{2^k} U_{kj}^\nu. \quad (17)$$

Clearly, this detector is obtained by combining three local adjacent bins in the wavelet space. It would be possible to combine two adjacent WT samples at the same scale, but it has proven to be less effective than hoped, and we do not report it here.

2) *Self-Normalizing Case*: For each column time-domain vector \mathbf{x}_i , let the $C_{k,j,i}$ s be the corresponding magnitude-squared DWT coefficients for each scale index $k = 0, 1, \dots, K-1$, intra-block time index $j = 1, \dots, 2^k$, and block index $i = 1, 2, \dots, L$. Similar to the frequency domain, we have

$$T_{wc} = \sum_{k=0}^{K-1} \sum_{j=1}^{2^k} z_{kj}^\nu \quad (18)$$

where

$$z_{kj} = \frac{C_{k,j,L}}{\frac{1}{L-1} \sum_{i=1}^{L-1} C_{k,j,i}}.$$

We record $U_{kji} = C_{k,j,i} + C_{k+1,2j-1,i} + C_{k+1,2j,i}$, for $k = 0, 1, \dots, K-1$, $j = 1, \dots, 2^k$, and $i = 1, 2, \dots, L$. As in the frequency domain, this combining approach suggests a new CFAR power-law detector

$$T_{wc3}(\mathbf{U}) = \sum_{k=0}^{K-1} \sum_{j=1}^{2^k} \left(\frac{U_{k,j,L}}{\frac{1}{L-1} \sum_{i=1}^{L-1} U_{k,j,i}} \right)^\nu \quad (19)$$

in the wavelet domain.

IV. PERFORMANCE ANALYSIS

Since we are interested in transients with unknown structure, location, and strength, our performance analysis will concentrate on the prediction of the threshold exceedance probability of the statistics under \mathbf{H}_0 and, thus, to choose a proper threshold to ensure a certain false alarm rate. We use the central limit theorem (CLT) to get the normal approximation to the distribution of the statistics. However, since the normal approximation provides poor approximation to deep tail probabilities, a procedure using saddle-point approximation is also introduced. The performance of different statistics are also studied via numerical simulation, where we set the total number of bins $N = 256$.

In the following analysis, we only consider detectors in the frequency domain; analysis in the wavelet domain is precisely equivalent, provided the transform is orthogonal.

A. Normal Approximation

• The Statistic T_{fc} :

This is a summation over iid random variables and, thus, must converge to a normal distribution by the central limit theorem [18]. Recall

$$T_{fc}(\mathbf{X}) = \sum_{j=1}^N \left(\frac{X_{jL}}{\frac{1}{L-1} \sum_{i=1}^{L-1} X_{ji}} \right)^\nu = \sum_{j=1}^N z_j^\nu \quad (20)$$

which converges in law to a normal distribution by CLT, provided $u_j = z_j^\nu$ has finite second moment. As shown before, z_j follows an iid $F(2, 2(L-1))$ distribution under \mathbf{H}_0 , and thus, we can compute the mean and variance of z_j^ν as

$$\begin{aligned} E(z_j^\nu) &= (L-1)^{\nu+1} \beta(L-\nu-1, \nu+1) \\ \text{Var}(z_j^\nu) &= (L-1)^{2\nu+1} [\beta(L-2\nu-1, 2\nu+1) \\ &\quad - (L-1) \beta(L-\nu-1, \nu+1)^2]. \end{aligned} \quad (21)$$

Having obtained this mean and variance, we have the following result for T_{fc} : Under the assumption that z_j s are

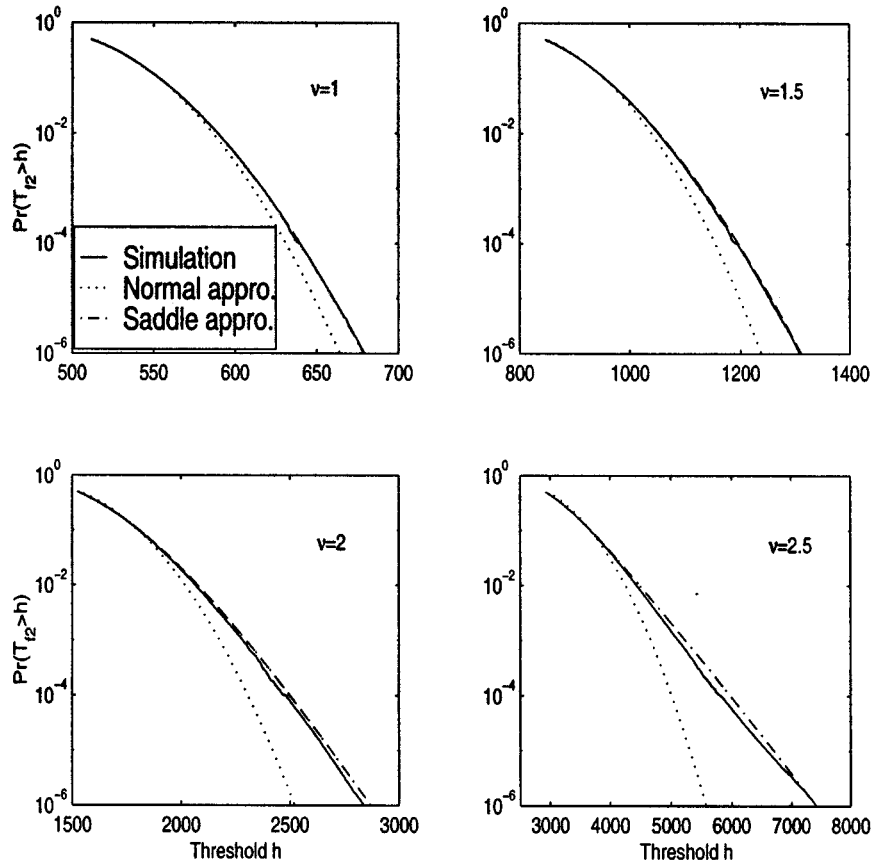


Fig. 5. Exceedance probability of T_{f2} with normal and saddlepoint approximations as a function of threshold h . The right-truncated distribution is used in the implementation of saddlepoint approximation. Here, $N = 256$, and $T = 20^\nu$. The approximation results are compared with the simulation results represented by the solid line.

iid $F(2, 2(L-1))$ for $j = 1, \dots, N$, T_{fc} converges in law for large N to $\mathcal{N}(N \cdot E(z^\nu), N \cdot \text{Var}(z^\nu))$.

- *The Statistic T_{f2} :*

We cannot use the classical CLT directly here since the detector is a summation over dependent, although identical, random variables. However, in the following, we show that the detectors converge to normal distributions using CLT after a rewriting of the statistics. Observe that we can rewrite T_{f2} as

$$\begin{aligned} T_{f2}(\mathbf{U}) &= \sum_{j=1}^N (X_{j-1} + X_j)^\nu = \sum_{j=1}^N U_j \\ &= \sum_{k=1}^{N/2} U_{2k-1} + \sum_{k=1}^{N/2} U_{2k} \equiv T_o + T_e \end{aligned} \quad (22)$$

where $U_j = (X_{j-1} + X_j)^\nu$. Since X_j are iid $\exp(\mu_0)$ under \mathbf{H}_0 (we assume $\mu_0 = 1$ for convenience), we know that U_{2k-1} , $k = 1, \dots, N/2$ are iid. Thus, T_o and T_e converge in law to $\mathcal{N}((N/2)\Gamma(\nu+2), (N/2)(\Gamma(2\nu+2) - \Gamma(\nu+2)^2))$ [recorded as $\mathcal{N}(\mu_2, \sigma_2^2)$] via the CLT.

Now, since $T_{f2} = T_o + T_e$, we know that T_{f2} follows distribution $\mathcal{N}(2\mu_2, \sigma_2^2(2+2\rho))$, where

$$\rho = \frac{E(T_o T_e) - E(T_o)E(T_e)}{\sqrt{\text{Var}(T_o)\text{Var}(T_e)}}$$

is the correlation coefficient, which is affected by the power law ν . If $\nu = 1$, we note that $T_1 = T_2$; thus, clearly, $\rho = 1$; if $\nu = 2$, our calculation reveals that $\rho = (20/21)$. For noninteger ν , a numerical method has to be used to calculate ρ . Fortunately, ρ is close to unity; thus, to simplify, we set $\rho = 1$ and thus approximate T_{f2} as convergent in law to $\mathcal{N}(2\mu_2, 4\sigma_2^2)$.

- *The Statistic T_{f3} :*

Approximating $\rho = 1$ as in the analysis of T_{f2} , we have that T_{f3} converges in law to $\mathcal{N}(3\mu_3, 9\sigma_3^2)$, where $\mu_3 = (N/3)\Gamma(\nu+3)/2$ and

$$\sigma_3^2 = \frac{N}{3} \left(\frac{\Gamma(2\nu+3)}{2} - \left(\frac{\Gamma(\nu+3)}{2} \right)^2 \right).$$

- *The Statistic T_{fc2} :*

For the detector T_{fc2} as defined in (15), we know

$$\left(\frac{U_{jL}}{\frac{1}{L-1} \sum_{i=1}^{L-1} U_{ji}} \right)$$

follows the $F(4, 4(L-1))$ distribution. Again, assuming that $\rho = 1$, we approximate T_{fc2} as convergent in law

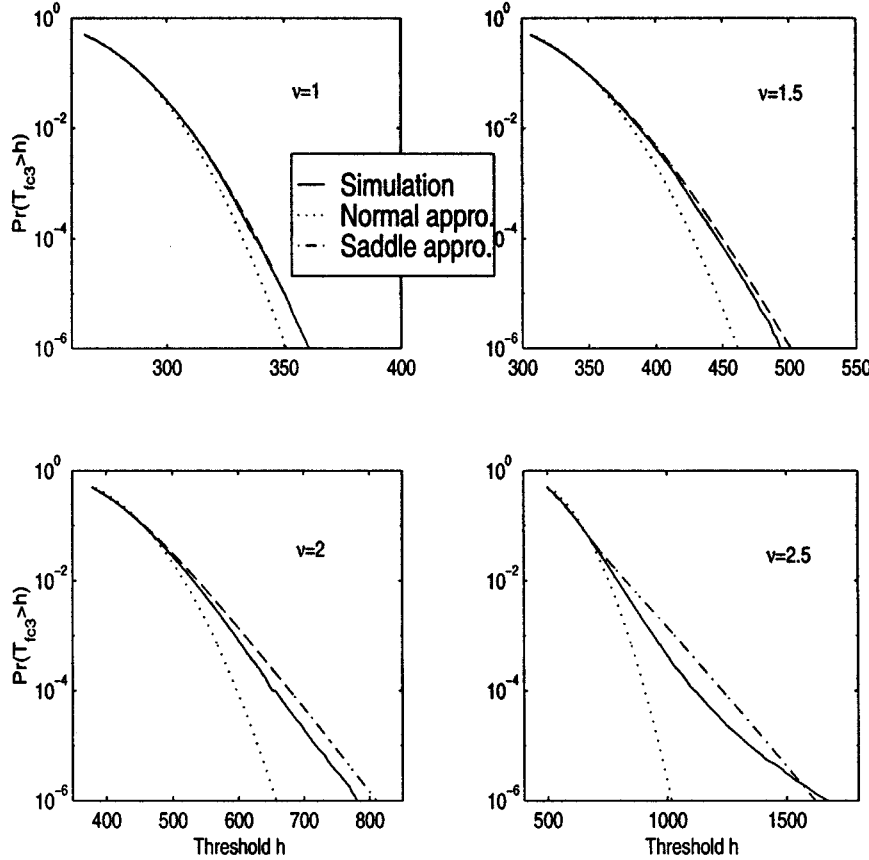


Fig. 6. Exceedance probability of T_{fc3} with normal and saddlepoint approximation as a function of threshold h . Here, $N = 256$, and $L = 10$. The right-truncated distribution is used in the implementation of saddlepoint approximation, and $T = 10^\nu$. The approximation results are compared with the simulation results represented by the solid line.

to $\mathcal{N}(2\mu_{c2}, 4\sigma_{c2}^2)$, where $\mu_{c2} = (N/2)E(Y)$ and $\sigma_{c2}^2 = (N/2)\text{Var}(Y)$

$$E(Y_j) = (L-1)^\nu \frac{\Gamma(2+\nu)\Gamma(2(L-1)-\nu)}{\Gamma(2(L-1))}$$

$$\text{Var}(Y_j) = (L-1)^{2\nu} \frac{\Gamma(2+2\nu)\Gamma(2(L-1)-2\nu)}{\Gamma(2)\Gamma(2(L-1))} - E(Y_j)^2. \quad (23)$$

• *The Statistic T_{fc3} :*

As in the previous cases, and using

$$\mu_{c3} = \frac{N}{3}(L-1)^\nu \frac{\Gamma(3+\nu)\Gamma(3(L-1)-\nu)}{\Gamma(3)\Gamma(3(L-1))}$$

$$\sigma_{c3}^2 = \frac{N}{3}(L-1)^{2\nu} \left(\frac{\Gamma(3+2\nu)\Gamma(3(L-1)-2\nu)}{\Gamma(3)\Gamma(3(L-1))} - \left(\frac{\Gamma(3+\nu)\Gamma(3(L-1)-\nu)}{\Gamma(3)\Gamma(3(L-1))} \right)^2 \right) \quad (24)$$

we approximate T_{fc3} as convergent in law to $\mathcal{N}(3\mu_{c3}, 9\sigma_{c3}^2)$.

B. Saddle-Point Approximation

The asymptotic normal distribution derived in the previous section is easy to work with. However, as we will see later, the

normal approximation tends to estimate the tail probability relatively poorly. To obtain more accurate performance evaluation, here, we introduce the saddle-point approximation method that can be thought as a refinement of normal approximation via the indirect use of the Edgeworth expansion. Here, we state the final result, and omit the detailed development; see [7].

Let X_1, \dots, X_n be iid with pdf $f(x)$, and let $\varphi(\theta) = E(e^{\theta X})$ be the corresponding Laplace transform defined for θ . We define the sample mean $X = (1/n)\sum_{i=1}^n X_i$ according to [7, eq. (2.2.6)], and hence, the tail probability can be approximated via the formula

$$\Pr(\bar{X} > x) = \frac{\varphi(\theta)^n e^{-n\theta x}}{\sqrt{n}|\theta|\sigma(\theta)} \times \left\{ B_0(\lambda) + \frac{\text{sgn}(\theta)}{\sqrt{n}} \frac{\zeta_3(\theta)}{6} B_3(\lambda) + \frac{1}{n} \left[\frac{\zeta_4(\theta)}{24} B_4(\lambda) + \frac{\zeta_3(\theta)^2}{72} B_6(\lambda) \right] + O(B_0(\lambda)n^{-3/2}) \right\} \quad (25)$$

where θ is the ML estimate of θ given x , $\lambda = \sqrt{n}|\theta|\sigma(\theta)$, $B_j(\lambda)$ is the Esscher function of j th order, and $\zeta_j(\theta)$ is the j th normal-

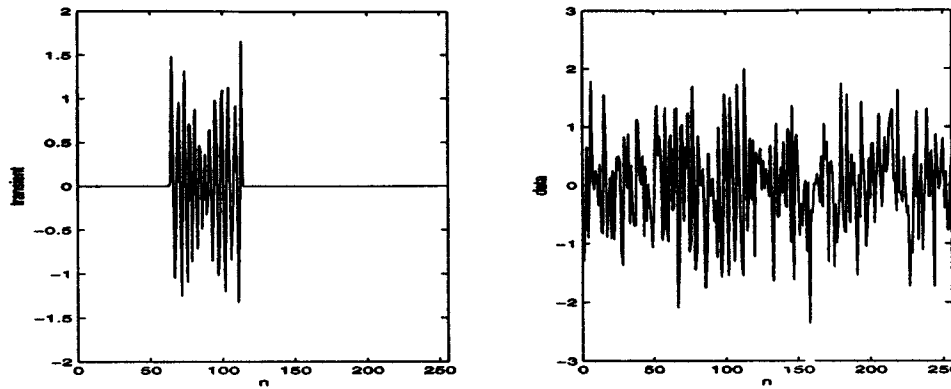


Fig. 7. Example of signal and observation process for Figs. 8 and 9. The signal (left panel) is created by passing white Gaussian noise through an FIR filter with a passband $0.4\pi < \omega < 0.6\pi$ (the number of signal-present FFT bins is approximately 25). On the right, noise is added.

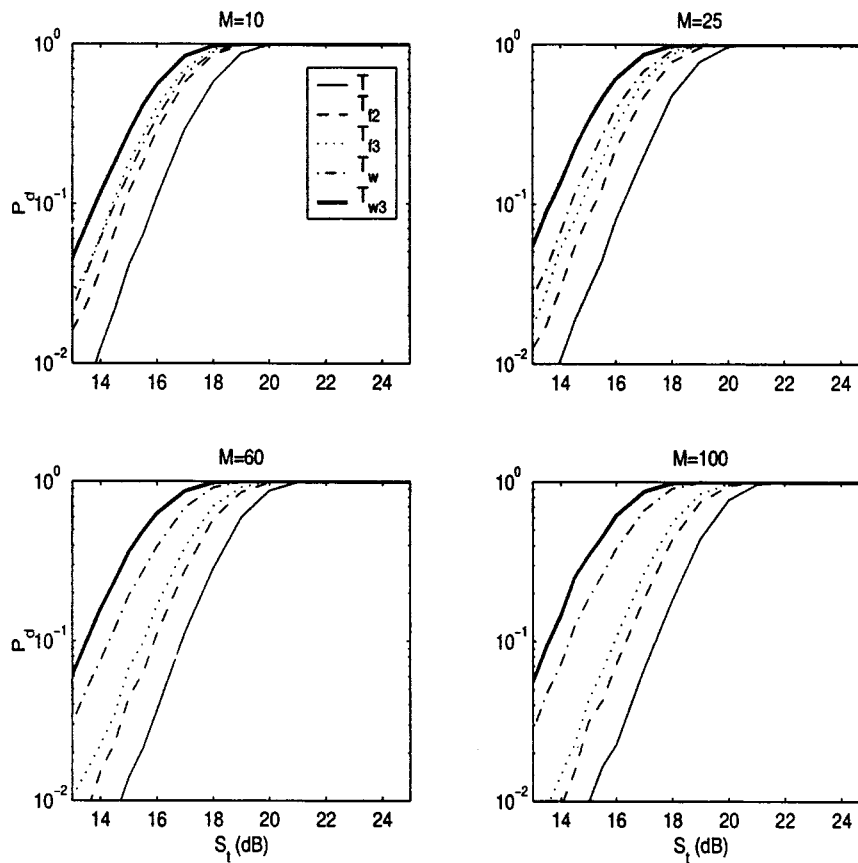


Fig. 8. Detection performances of new power-law statistics in the frequency and the wavelet domains in the prenormalized case. The exponent in each case is $\nu = 2.5$, and the transient duration is $M_t = 20$ samples; different panels refer to the number of frequency bins occupied by the signal.

ized cumulant. A detailed description of the above quantities is available in [7].⁴

C. Comparison of Approximations with Simulation

As mentioned earlier, the performance analysis in this section is focused on the exceedance probability of the statistics under the \mathbf{H}_0 hypothesis. The results obtained using the normal approximation and the saddlepoint approximation are shown and compared with results of numerical simulations based on

⁴In our case, for X_i obeying pdf $f(x)$, the moment generation function (MGF) does not exist. To resolve this, we simply truncate X_i to a value T .

10^8 runs, for different statistics and different exponents ν , with $N = 256$.

For the white noise (prenormalized) case, we investigate the statistics T_{f2} and T_{f3} that combine two and three contiguous FFT bins correspondingly and show the performance analysis of T_{f2} in Fig. 5. We can see that the normal approximation tends to underestimate the tail probability, whereas the saddlepoint approach shows good prediction. It implies both that saddlepoint approximation is an efficient method to analyze the performances of our statistics T_{f2} and T_{f3} and that setting the correlation coefficient ρ as 1 is an acceptable approximation.

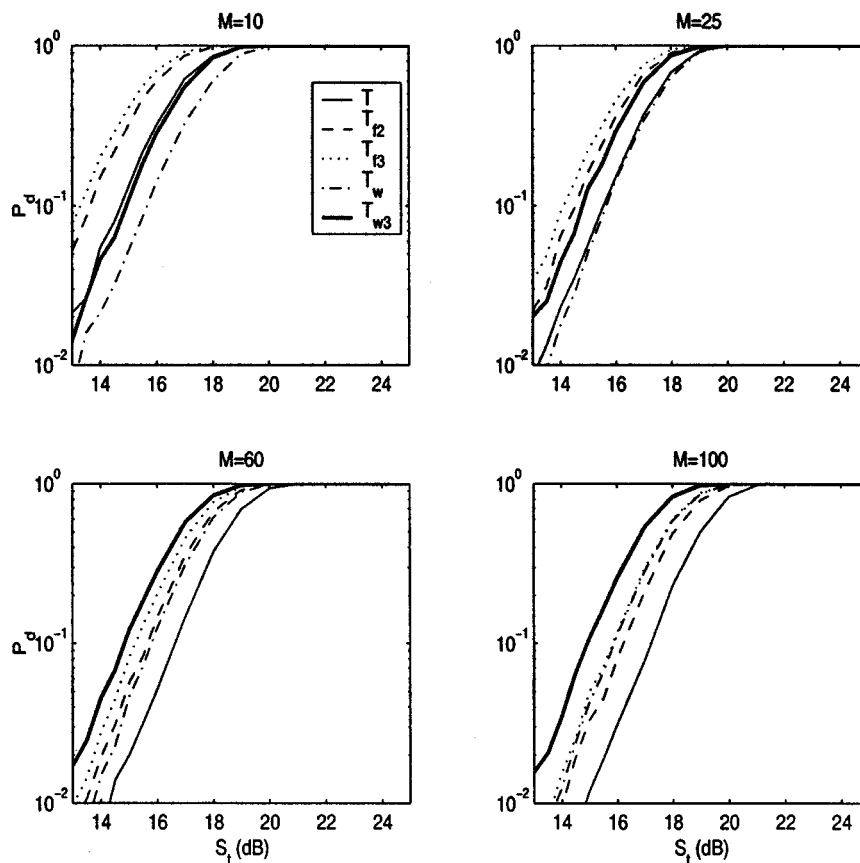


Fig. 9. Detection performances of new power-law statistics in the frequency and the wavelet domains in the prenormalized case. The exponent in each case is $\nu = 2.5$, and the transient duration is $M_t = 50$ samples; different panels refer to the number of frequency bins occupied by the signal.

For the colored noise (self-normalizing) case, we investigate the statistics T_{fc} , T_{fc2} , and T_{fc3} and show the result of T_{fc3} in Fig. 6. According to our earlier SNR analysis, we are interested in $1 < \nu < 2$. Here, we set $N = 256$ and $L = 10$ to be consistent with our later P_d simulations. We can see that the normal approximation estimates the tail probability rather poorly, whereas the saddlepoint approach shows much better prediction. It is noted, however, that even the saddlepoint approximation tends to mismatch the tail probability as ν grows large.

V. PERFORMANCE COMPARISON

Here, we apply the detectors developed in the previous sections to numerical examples. For fixed $P_{fa} = 10^{-4}$, applying the thresholds obtained in Section IV, we compare probabilities of detection against aggregate SNR. The *aggregate* SNR is the total signal energy divided by the total noise variance, meaning that the *per-sample* (over the entire block and not just for the fuzzily-defined duration of the transient signal) SNR should be divided by this number; for example, excellent performance is available at an aggregate SNR of 20 dB, and this translates to -4 dB per sample.

Pre-normalized Data: The detection performance of the improved detectors in the frequency and the wavelet domains are compared with the power law of [12] with exponent $\nu = 2.5$. Examples of the signal and noise are shown in Fig. 7 for a number of signal-containing frequency bins $M \approx 25$. From Fig. 8, in which the time-domain transient signal is of length

$M_t = 20$ samples, it is clear that combining two or three contiguous FFT or wavelet bins together does improve the detection performance over different SNRs S_t . It is also noted that the detector T_{w3} based on the contiguity of wavelet bins shows advantages over all others, and the explanation is presumably that T_{w3} utilizes both temporal and frequency contiguity. However, in the case of a longer transient signal (see Fig. 9) for which the length is $M_t = 50$ samples, those transients that are more concentrated in the frequency domain ($M = 10$ and $M = 25$) are best detected by the FFT-based statistic T_{f3} . This is further explored in Fig. 10; here, contours of the probability of detection are plotted on transient-length (vertical) and bandwidth (horizontal) axes for detectors T_w , w_3 , T_f , and T_{f3} . It is clear that the wavelet-based detectors are more forgiving than those based on the short-time frequency transformation, but that transients of sufficiently narrow bandwidth and broad length are best served in the latter domain.

Self-Normalizing Case: The results of detectors in the case that self-normalization is required are shown in Fig. 11. From comparison with Fig. 9 (the prenormalized case), it is clear that the losses arising from the need to normalize are relatively minor. It is also gratifying that the statistic T_{wc3} is, in all these cases, the best.

VI. SUMMARY

In [12], Nuttall derived and justified a new and easy-to-implement statistic for the detection of short-duration (transient)

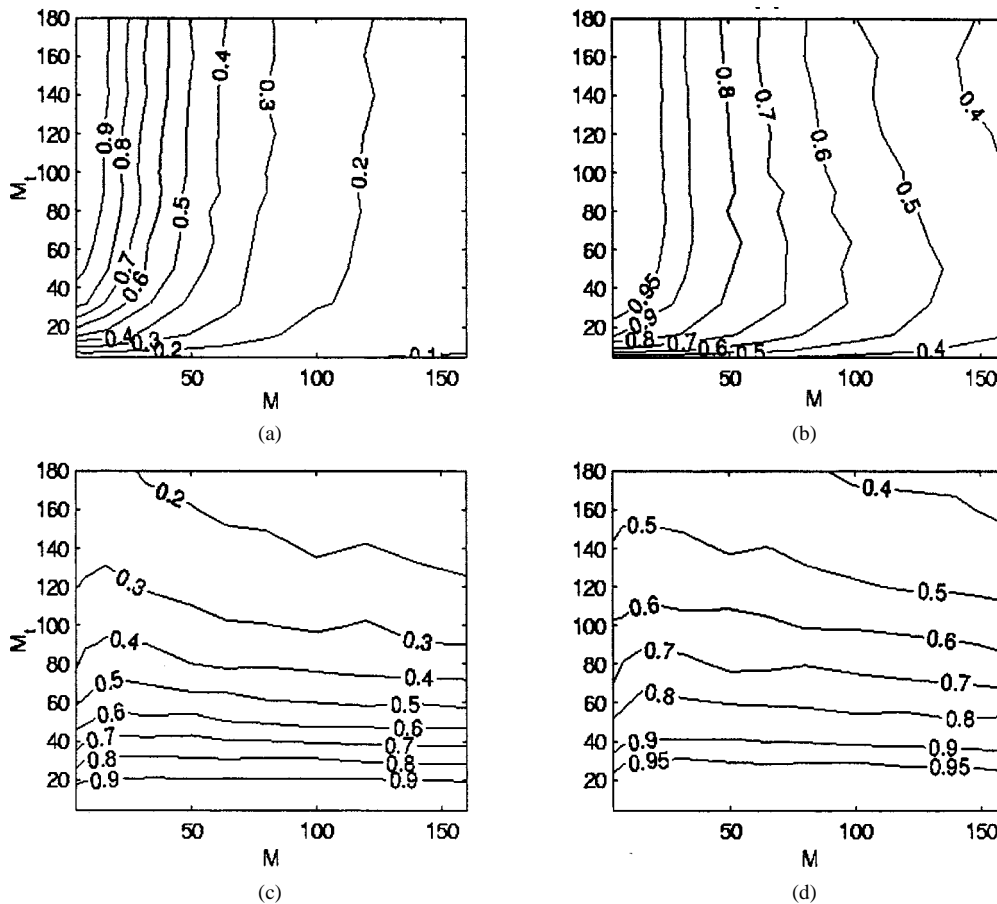


Fig. 10. Probability of detection contours for (a) T_f , (b) T_{f3} , (c) T_w , and (d) T_{w3} . These are plotted versus transient length M_t and number of signal-occupant frequency bins M .

signals: the sum of magnitude-square DFT outputs from a block of N time domain data, each raised to a power typically in the range of 2 to 3. This test has been found to be very effective indeed.

The power-law detector is almost a plug-in transient detector for all purposes but not quite: Prewhitened and prenormalized data is required. We have thus extended the power-law detector to be self-normalizing by raising to an exponent not the DFT data directly but, instead, the power in each DFT bin relative to the average power in previous DFTs. While Nuttall provided some justification both for the exponentiation and exponent in the original power-law test, their applicability in the self-normalizing case is not straightforward. Consequently, a mode of analysis (“input SNR loss”) has been proposed for the choice of exponent. It is found that the optimal exponent is somewhat lower than in the original (prenormalized) power-law case.

At any rate, the self-normalizing solution works very nicely. All that is needed to have an all-purpose transient detector with simple implementation is some means to set the threshold, and this is provided via a saddle-point approximation.

Along the path to development of the improved CFAR power-law detector, it was noted that there is a tendency among real transient signals for energy to aggregate in nearby DFT bins (i.e., to be bandlimited to some degree). In their formulations, neither the original nor CFAR power-law detectors take advantage of this, and consequently, a combined-bin power-law detector is proposed. In experiments, versions of

this are shown to offer significant improvement. These are made self-normalizing, and a saddlepoint approximation for threshold setting is provided. It was additionally noted that the power-law dogma of preprocessing via the DFT is open to the challenge, and indeed, a power-law processor operating on (Haar) wavelets is developed, made self-normalizing, augmented to use combined bins (since transient signals most transient signals are aggregated not just in frequency but in time/scale as well), and accorded a saddlepoint approximation for threshold setting.

Quite a few of the detectors have been developed and analyzed in this paper. We note that beyond easy choices such as window size (for CFAR) or type (wavelet/DFT or agglomerating/single-bin) and guided selection of the exponent, there are no parameters to tune. We thus consider them to be “plug-in” transient detectors, and since they make minimal assumptions on the structure of the transient signal, save that it have some degree of concentration of energy in time and/or frequency, we advertise them as “all-purpose.” For reference, we give their taxonomy in Table I.

Note that the choice of prenormalized versus self-normalizing depends on the data, but in either case, our overall conclusion is that although all of these tests work well, the combined/wavelet power-law detectors (if data are prenormalized, T_{w3} , and if self-whitening and CFAR is necessary, T_{wc3}) are perhaps the finest of all. The statistics are compellingly simple to use. Take a multiresolution decomposition using the Haar

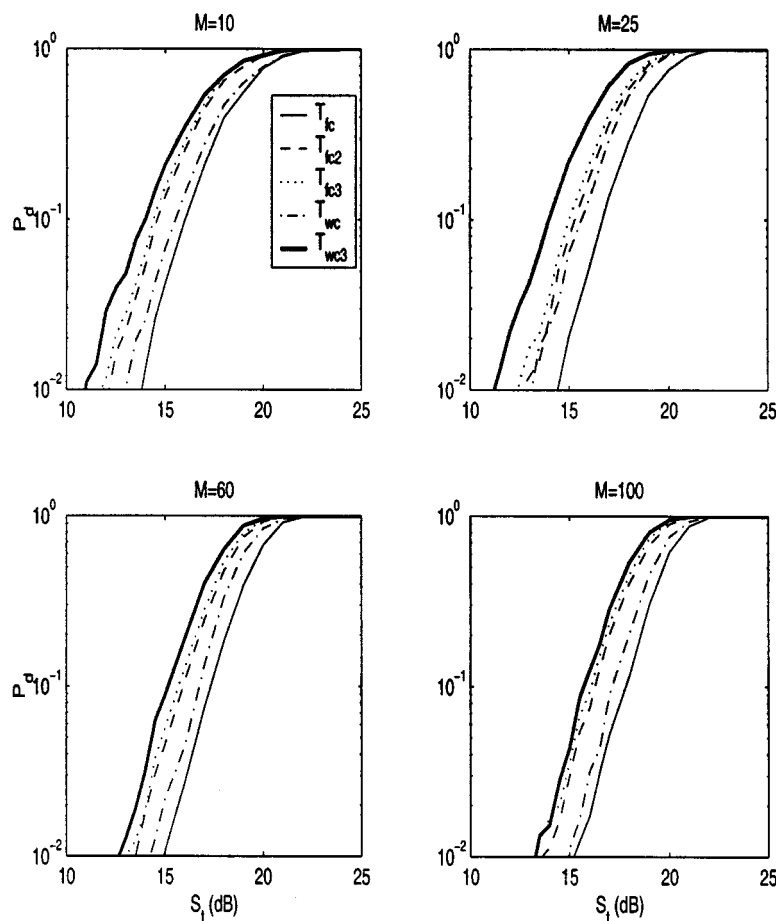


Fig. 11. Detection performance of power-law detectors in the frequency and the wavelet domains for transient detection in colored noise. Here, the time-domain transient length is $M_t = 50$, and the exponent used in all cases is $\nu = 1.5$.

TABLE I

CATEGORIZATION OF VARIOUS TRANSIENT DETECTORS DISCUSSED IN THIS PAPER. NUTTALL'S CFAR POWER-LAW IS "PARTIALLY" SELF-NORMALIZING IN THAT IT IS COMPLETELY INSENSITIVE TO SCALE BUT HAS NO MEANS TO DEAL WITH NONWHITE DATA

detector	provenance	pre-processing	bins combined	self-normalizing
T	Nuttall	DFT	1	no
T_{cpt}	Nuttall	DFT	1	partial
T_{f2}	new	DFT	2	no
T_{f3}	new	DFT	3	no
T_{fc}	new	DFT	1	yes
T_{fc2}	new	DFT	2	yes
T_{fc3}	new	DFT	3	yes
T_w	new	wavelet	1	no
T_{w3}	new	wavelet	3	no
T_{wc}	new	wavelet	1	yes
T_{wc3}	new	wavelet	3	yes

basis, normalize the magnitude-square by previous values at each scale (in the self-whitening case only), combine the result in groups of three according to Fig. 4, exponentiate (to a power 2.5 if prenormalized and to the power 1.5 if normalized), and sum. The resulting statistic can be relied on to detect quite a wide range of transient signals below (often considerably below) -3 dB on a sample-by-sample basis.

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